# Lecture 5: Spectral Analysis and Decomposition Introduction to Time Series, Fall 2023 Ryan Tibshirani

Related reading: Chapters 4.1–4.3 of Shumway and Stoffer (SS).

### 1 Periodic processes

• Consider a periodic process of the form

$$x_t = A\cos(2\pi\omega t + \phi) \tag{1}$$

• Importantly, the quantity  $\omega$  in the above definition is called *frequency* of the process; and the quantity  $1/\omega$  is called the *period* or *cycle*. As t varies from 0 to  $1/\omega$ , note that the process goes through one complete cycle (it ends up back where it started). See Figure 1

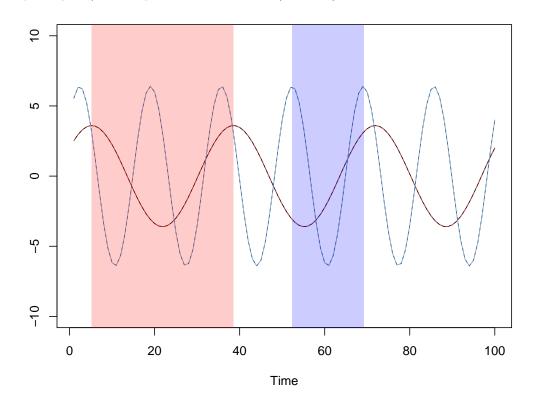


Figure 1: Two examples of cosine processes, the first (in red) having a frequency  $\omega = 3/100$  and amplitude  $\sqrt{2^2 + 3^2} \approx 3.6$ , and the second (in blue) having a frequency  $\omega = 6/100$  and amplitude  $\sqrt{4^2 + 5^2} \approx 6.4$ .

- The quantity A is called the *amplitude* and  $\phi$  the *phase* of the process. The amplitude controls how high the peaks are, and the phase determine where (along the cosine cycle) the process starts at the origin t = 0
- We can introduce randomness into the process (1) by allowing A and  $\phi$  to be random

• It will be useful to reparametrize. In general, recall the trigonometric identity (cosine compound angle formula):

$$\cos(a+b) = \cos(a)\cos(b) - \sin(a)\sin(b) \tag{2}$$

Thus, starting with (1), we can rewrite this as  $x_t = A\cos(\phi)\cos(2\pi\omega t) - A\sin(\phi)\sin(2\pi\omega t)$ . Simply letting  $U_1 = A\cos(\phi)$ ,  $U_2 = -A\sin(\phi)$ , we can therefore write

$$x_t = U_1 \cos(2\pi\omega t) + U_2 \sin(2\pi\omega t) \tag{3}$$

with  $U_1, U_2$  our two random variables, determining the amplitude of the cosine and sine components separately

• Note that another way of writing the relationship between  $A, \phi$  and  $U_1, U_2$  is (why?):

$$A = \sqrt{U_1^2 + U_2^2}, \quad \phi = \tan^{-1}(-U_2/U_1)$$

• An interesting fact (that you can try to verify as a challenge):

$$U_1, U_2 \sim N(0, 1)$$
, independently  $\iff A \sim \chi_2^2, \phi \sim \text{Unif}(-\pi, \pi)$ , independently

#### 1.1 Stationarity

- If  $U_1, U_2$  are uncorrelated, each with mean zero and variance  $\sigma^2$ , then the periodic process  $x_t, t = 1, 2, 3, \ldots$  defined in (3) is stationary
- To check this: simply compute the mean function

$$\mu_t = \mathbb{E}(x_t) = 0$$

which is constant in time; and the auto-covariance function

$$\begin{aligned} \gamma(s,t) &= \operatorname{Cov}(x_s, x_t) \\ &= \operatorname{Cov}\left(U_1 \cos(2\pi\omega s) + U_2 \sin(2\pi\omega s), U_1 \cos(2\pi\omega t) + U_2 \sin(2\pi\omega t)\right) \\ &= \operatorname{Cov}\left(U_1 \cos(2\pi\omega s), U_1 \cos(2\pi\omega t)\right) + \operatorname{Cov}\left(U_2 \sin(2\pi\omega s), U_1 \cos(2\pi\omega t)\right) \\ &+ \operatorname{Cov}\left(U_1 \cos(2\pi\omega s), U_2 \sin(2\pi\omega t)\right) + \operatorname{Cov}\left(U_2 \sin(2\pi\omega s), U_2 \sin(2\pi\omega t)\right) \\ &= \sigma^2 \cos(2\pi\omega s) \cos(2\pi\omega t) + 0 + 0 + \sigma^2 \sin(2\pi\omega s) \sin(2\pi\omega t) \\ &= \sigma^2 \cos(2\pi\omega (s-t)) \end{aligned}$$

which only depends on the lag s - t (where in the last line we used the identity (2) once again)

### **1.2** General mixtures

• As a generalization of (3), we can also mix together a total of p periodic processes, defining

$$x_{t} = \sum_{i=1}^{p} \left( U_{j1} \cos(2\pi\omega_{j}t) + U_{j2} \sin(2\pi\omega_{j}t) \right)$$
(4)

for  $U_{j1}, U_{j2}, j = 1, \ldots, p$  all uncorrelated random variables with mean zero, where  $U_{j1}, U_{j2}$  have variance  $\sigma_j^2$ 

• As a generalization of the above calculation, you'll show on your homework that the process  $x_t$ ,  $t = 1, 2, 3, \ldots$  defined in (4) is stationary, with auto-covariance function

$$\gamma(h) = \sum_{j=1}^{p} \sigma_j^2 \cos(2\pi\omega_j h)$$

• Figure 2 displays a couple of mixture processes of the form (4) (with p = 2 and p = 3). Note the regular repeating nature of the mixture processes. One might wonder how we can decompose a such a mixture into its frequency components (periodic processes, each of the form (3)). This is, in fact, one of the main objectives in spectral analysis

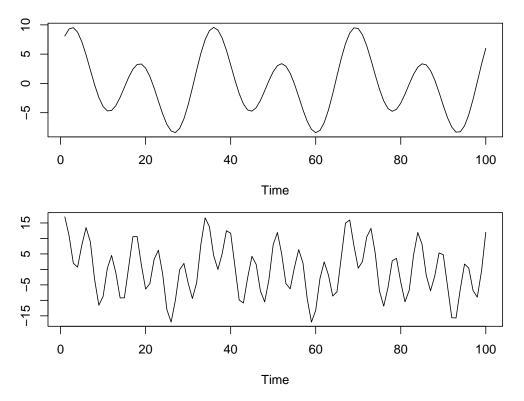


Figure 2: Mixture of periodic processes of different frequencies (and amplitudes).

• And the answer, as we'll see next, is given by something you're already quite familiar with ... regression!

## 2 Fourier decomposition

• Given a time series  $x_t$ , t = 1, ..., n, consider *seeking* a decomposition like (4). We could do this by regressing this time series onto cosine and sine features of different frequencies,

$$c_{tj} = \cos(2\pi j/n \cdot t), \quad t = 1, \dots, n$$
  
$$s_{tj} = \sin(2\pi j/n \cdot t), \quad t = 1, \dots, n$$

(which we call these "basis functions" in the context of this particular regression problem), and so the regression model is

$$x_t \approx a_0 + \sum_{j=1}^p (a_j c_{tj} + b_j s_{tj}), \quad t = 1, \dots, n$$
 (5)

Hence the regression coefficients  $a_j, b_j, j = 1, ..., p$  represent the amplitudes

• How large should p in the above regression model? That is, how many cosine and sine basis functions do we need? An amazing fact (at least, it will probably seem amazing if you've never seen Fourier decomposition before): for any time series  $x_t$ , t = 1, ..., n, we only need to set p = (n-1)/2, and then the representation in (5) will be exact!

- That is, there are coefficients  $\hat{a}_j, \hat{b}_j, j = 1, \dots, p$  that will give us an equalities in (5), for all t
- (This assumes that n is odd; if n is even, then we need to add an additional component  $a_{n/2} \cos(\pi t) = a_{n/2}(-1)^t$ , and the same claim holds: the representation is exact)
- To find the coefficients  $\hat{a}_j, \hat{b}_j, j = 1, ..., p$ , we can simply perform regression (least squares). We let  $x \in \mathbb{R}^n$  denote our time series represented as a vector, which serves as the response vector in our regression problem, and we assemble our cosine and sine basis functions into a feature matrix

$$Z = \begin{bmatrix} \frac{1}{\sqrt{2n}} & \cos(2\pi\frac{1}{n}\cdot1) & \sin(2\pi\frac{1}{n}\cdot1) & \cos(2\pi\frac{2}{n}\cdot1) & \dots & \sin(2\pi\frac{n-1}{2n}\cdot1) \\ \frac{1}{\sqrt{2n}} & \cos(2\pi\frac{1}{n}\cdot2) & \sin(2\pi\frac{1}{n}\cdot2) & \cos(2\pi\frac{2}{n}\cdot2) & \dots & \sin(2\pi\frac{n-1}{2n}\cdot2) \\ \vdots & & & \\ \frac{1}{\sqrt{2n}} & \cos(2\pi\frac{1}{n}\cdotn) & \sin(2\pi\frac{1}{n}\cdotn) & \cos(2\pi\frac{2}{n}\cdotn) & \dots & \sin(2\pi\frac{n-1}{2n}\cdotn) \end{bmatrix} \in \mathbb{R}^{n \times n}$$

• We can then perform regression of x on Z in order to estimate the coefficients:

$$(Z^{\mathsf{T}}Z)^{-1}Z^{\mathsf{T}}x$$

- However, something is very special about out matrix Z: it satisfies  $Z^{\mathsf{T}}Z = (n/2)I$ , where I the  $n \times n$  identity matrix. In other words, *its columns are uncorrelated, and have squared*  $\ell_2$  *norm equal to* n/2. This is a very special property of the cosine and sine basis functions (and it is the foundation of the discrete Fourier transform, to be discussed shortly)
- Thus, writing  $z_j$ , j = 1, ..., n as the columns of Z, we have

$$(Z^{\mathsf{T}}Z)^{-1}Z^{\mathsf{T}}x = \begin{bmatrix} \frac{2}{n}z_1^{\mathsf{T}}x\\ \frac{2}{n}z_2^{\mathsf{T}}x\\ \vdots\\ \frac{2}{n}z_n^{\mathsf{T}}x \end{bmatrix},$$

so the multiple regression coefficients of x on Z are simply the marginal regression coefficients

• More explicitly, the coefficients are  $\hat{a}_0 = \bar{x}$ , and

$$\hat{a}_{j} = \frac{2}{n} c_{j}^{\mathsf{T}} x = \frac{2}{n} \sum_{t=1}^{n} x_{t} \cos(2\pi j/n \cdot t)$$

$$\hat{b}_{j} = \frac{2}{n} s_{j}^{\mathsf{T}} x = \frac{2}{n} \sum_{t=1}^{n} x_{t} \sin(2\pi j/n \cdot t)$$
(6)

• And to be perfectly clear, this gives us the exact decomposition

$$x_t = \bar{x} + \sum_{j=1}^{(n-1)/2} \left( \hat{a}_j \cos(2\pi j/n \cdot t) + \hat{b}_j \sin(2\pi j/n \cdot t) \right), \quad t = 1, \dots, n$$
(7)

The reason: because Z is orthogonal, it has n linearly independent columns in n dimensions, so we can exactly represent any vector as a linear combination of its columns—which means that the fitted values from regressing x on Z will be exactly x

• (Side note on computation: at first glance, in order to compute each  $\hat{a}_j$  or  $\hat{b}_j$ , we require O(n) operations, and so computing all of them should take  $O(n^2)$  time ... but in fact the *entire set of* coefficients  $\hat{a}_j, \hat{b}_j, j = 1, ..., (n-1)/2$  can be computed in  $O(n \log n)$  time using what is known as the fast Fourier transform, which we'll return to below)

#### 2.1 Periodogram

• Given a series  $x_t$ , t = 1, ..., n, we can define an object from the coefficients (6) in the decomposition (7) that is called the *periodogram*, denoted  $P_x$ . This takes values at frequencies j/n, for j = 1, ..., (n - 1)/2, and is defined by

$$P_x(j/n) = \frac{n}{4} \left( \hat{a}_j^2 + \hat{b}_j^2 \right)$$
(8)

When the underlying series is clear from the context, we will drop the subscript and simply write P

• Large values of the periodogram indicate which frequencies are predominant in the given series. This is illustrated in Figure 3, which displays the periodogorams for the two series in Figure 2. Another nice real data example (from a 1923 texbook on numerical analysis!) is given in Figure 4

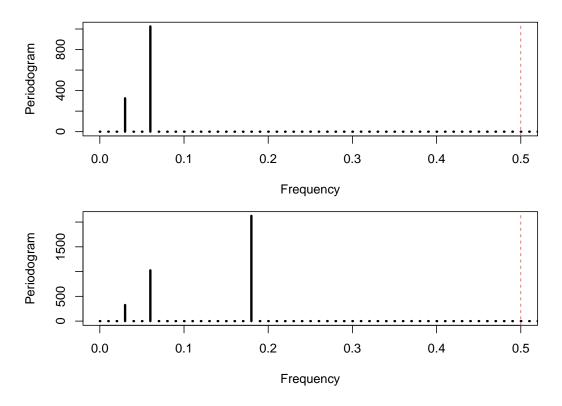


Figure 3: Periodograms for the mixture processes in Figure 2.

• If we think back to the mixture process (4) as a model for our data, then the periodogram gives us a breakdown of which frequencies are the largest sources of variance: recall  $\sigma_j^2 = \mathbb{E}(U_{j1}^2 + U_{j2}^2)/2$  is the variance at frequency  $\omega_j$ 

### 2.2 Discrete Fourier transform

- The coefficients (6) used in the decomposition (7) can be computed efficiently by recognizing their connection to what is called the *discrete Fourier transform* (DFT)
- The DFT of a series  $x_t$ , t = 1, ..., n, is denoted  $d_x$  and defined as

$$d_x(j/n) = \frac{1}{\sqrt{n}} \sum_{t=1}^n x_t \exp(-2\pi i j/n \cdot t), \quad j = 0, 1, \dots, n-1$$
(9)

where *i* is the imaginary unit, which satisfies  $i^2 = -1$ . Thus the DFT is complex-value. As before, when the underlying series is clear from the context, we will drop the subscript and simply write *d* 

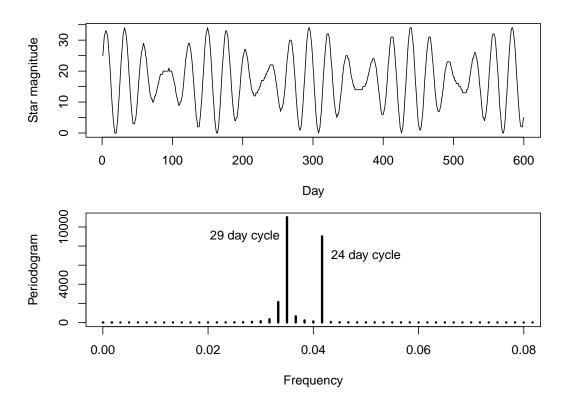


Figure 4: Periodogram for star magnitude data over 600 consecutive days (originally from "The Calculus of Observations" by Whittaker and Robinson, adapted by SS). Note the large values of the periodogram at 0.35 and 0.41, which correspond to  $1/0.35 \approx 29$  and  $1/0.41 \approx 24$  day cycles. For more on the interpretation, see Example 4.3 of SS.

• Recalling Euler's formula,  $e^{i\theta} = \cos(\theta) + i\sin(\theta)$ , we also have (using the fact that cosine is an even function and sine is odd):

$$d(j/n) = \frac{1}{\sqrt{n}} \sum_{t=1}^{n} x_t \cos(2\pi j/n \cdot t) - \frac{i}{\sqrt{n}} \sum_{t=1}^{n} x_t \sin(2\pi j/n \cdot t), \quad j = 0, 1, \dots, n-1$$

• Thus, from the DFT, we can compute each cosine and sine coefficient in (6) by

$$\hat{a}_j = \frac{2}{\sqrt{n}} \operatorname{Re}\{d(j/n)\}$$
 and  $\hat{b}_j = -\frac{2}{\sqrt{n}} \operatorname{Im}\{d(j/n)\}$ 

where for a complex number z = a + bi, we use  $\operatorname{Re}\{z\} = a$  and  $\operatorname{Im}\{z\} = b$  to denote its real and imaginary parts

• Note the following interesting connection to the periodogram. Since the the modulus of each entry of the DFT satisfies (by definition)  $|d(j/n)|^2 = \operatorname{Re}\{d(j/n)\}^2 + \operatorname{Im}\{d(j/n)\}^2$ , the periodogram in (8) is

$$P(j/n) = \frac{n}{4} (\hat{a}_j^2 + \hat{b}_j^2)$$
  
=  $\frac{n}{4} \left( \frac{4}{n} \operatorname{Re} \{ d(j/n) \}^2 + \frac{4}{n} \operatorname{Im} \{ d(j/n) \}^2 \right)$   
=  $|d(j/n)|^2$  (10)

• In other words, the periodogram is simply the squared modulus of the DFT!

- Side note: the entire DFT can be computed rapidly using an algorithm called the fast Fourier transform (FFT), which takes  $O(n \log n)$  operations (most efficient in practice when n is a highly composite integer, such as a power of 2). The modern generic FFT algorithm is credited to Cooley and Tukey in the 1960s, but similar ideas were arround much earlier
- Side side note: different software implementations scale the FFT/DFT differently, so you have to be careful to consult the documentation. For example, the fft() function in R computes it without the leading factor of  $n^{-1/2}$ , and with an additional factor of  $\exp(2\pi i j/n)$ , but this doesn't matter since we're only using it in our examples to compute the squared modulus, i.e., the periodogram, across frequencies

### 3 Interlude: ST decomposition

- As an interlude, we'll demo how to use the periodogram, in combination with smoothing (as we learned in the last lecture) to detect the presence of seasonality in time series, and fit a seasonal-trend (ST) decomposition
- Warning: what we do here is very simple and may very well offend researchers well-versed in classical time series decomposition and econometrics alike. It is only meant as a demo. We repeat: it is just a demo!
- (You can learn more about decomposition methods in time series in Chapters 3.4–3.6 of the HA book. It is worth mentioning that in statistics, the STL decomposition is pretty standard and popular, which stands for "seasonal-trend decomposition using LOESS", with LOESS being a particular type of smoother that we didn't cover, but it acts like kernel smoother with a varying bandwidth. For an econometrics perspective, see the references from the last lecture: Hodrick-Prescott → Hamilton → Hodrick again. The last reference is especially scholarly and reviews what has been done over the years)
- Ok with all those caveats laid out, a pretty simple and generic method to perform an ST decomposition of a time series  $y_t$ , t = 1, ..., n is as follows:
  - Use a smoother to estimate a trend  $\hat{\theta}_t$ , t = 1, ..., n, aiming to undersmooth somewhat so as to ignore a (possible) cyclic or seasonal component
  - Compute residuals  $r_t = y_t \hat{\theta}_t, t = 1, \dots, n$
  - Absent any prior knowledge about the periods of the seasonal component (i.e., without wanting to specify a priori that there might be weekly, monthly, quarterly, etc. components), just compute the periodogram of the residuals
  - Identify large peaks in the periodogram, and fit and add to the model seasonal components that correspond to the predominant periods (inverse frequencies) that are present
- Of course the method presented above is limited in several ways (e.g., it assumes that the seasonal components have constant amplitude over time), and other methods are more advanced in various ways. You can read more in the aforementioned references if you are curious
- Figure 5 gives an example applied to US retail employment data. We can see very clear cycles at about a year, half-year, third-year, and quarter-year

## 4 Spectral representation

- Now we turn to a general fact about stationary processes in time series, which is called the spectral representation of the auto-covariance function
- In particular, if  $x_t, t = 1, 2, 3, ...$  is stationary with auto-covariance function  $\gamma(h) = \text{Cov}(x_{t+h}, x_t)$ , then there exists a unique increasing function F, called the *spectral distribution function* correspond-

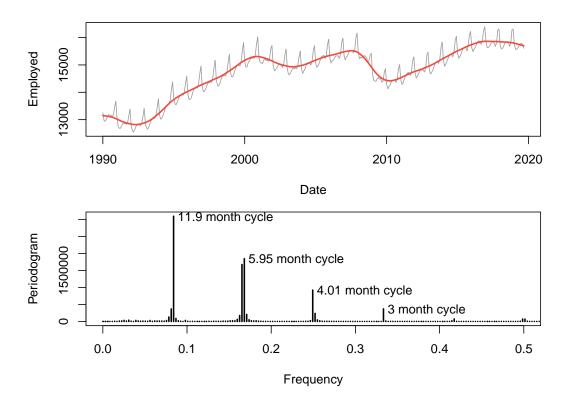


Figure 5: ST decomposition, using an HP filter for the smoother, and a periodogram on the residuals to detect predominant frequency components.

ing to the process, with  $F(\omega) = 0$  for  $\omega \leq -1/2$ , and  $F(\omega) = \gamma(0)$  for  $\omega \geq 1/2$ , such that for any  $h = 0, \pm 1, \pm 2, \ldots$ ,

$$\gamma(h) = \int_{-1/2}^{1/2} \exp(2\pi i\omega h) \, dF(\omega)$$

- Above, we can think of F as being analogous to a cumulative distribution function (CDF), and thus the integral as being analogous as an expectation defined with respect the distribution that governs F. The only difference is that the total mass here of need not be 1, but is instead  $\gamma(0)$
- We will generally ignore this distinction and call F a distribution anyway, and in the case F is differentiable, we will denote its derivative by f and call this the *spectral density*
- When the spectral density exists,<sup>1</sup> note that we have the representation

$$\gamma(h) = \int_{-1/2}^{1/2} \exp(2\pi i\omega h) f(\omega) d\omega, \quad h = 0, \pm 1, \pm 2, \dots$$

• There is also an inverse relationship: we can represent f in terms of  $\gamma$ , via

$$f(\omega) = \sum_{h=-\infty}^{\infty} \gamma(h) \exp(-2\pi i\omega h), \quad \omega \in [-1/2, 1/2]$$
(11)

In other words, the spectral density and auto-covariance function are Fourier transform pairs

<sup>&</sup>lt;sup>1</sup>It exists when the auto-covariance function is absolutely summable, which means that it satisfies  $\sum_{h=-\infty}^{\infty} |\gamma(h)| < \infty$ . See Appendix C of SS for details.

- The important high-level perspective to remember here: the auto-covariance function and spectral density contain the same information about a time series, but express it in different ways
- The auto-covariance function expresses the variation broken down by *lags*, whereas the spectral density expresses variation broken down by *frequencies*—or by *cycles* (remembering that the inverse of a frequency is a cycle)
- Next we compute the spectral density in a number of our favorite example stationary time series models. It will be helpful to point out that  $\gamma(h) = \gamma(-h)$  implies  $f(\omega) = f(-\omega)$ , so we only need to keep track of f for  $\omega \in [0, 1/2]$

### 4.1 White noise

• Recall our most basic stationary process which is white noise:  $x_t, t = 1, 2, 3, ...$  are uncorrelated random variables, with zero mean, and constant variance  $\sigma^2$ . So

$$\gamma(h) = \begin{cases} \sigma^2 & h = 0\\ 0 & h \neq 0 \end{cases}$$

• The spectral density, according to (11), is therefore

$$f(\omega) = \sigma^2, \quad \omega \in [0, 1/2]$$

- So, white noise has the simplest spectral density there is: it is constant!
- For the moment you (may) have been waiting for: we can finally explain where a lot of the nomenclature is coming from ... if we think about a time series as being comprised of a mixture of colors which we can generally do since any time series has the periodic representation (7)—then *spectral analysis* provides us a tool like a prism, for decomposing this series into its primary colors, or *spectra*. And, just like the color white, a *white noise* series is an equal mix of all colors (frequencies)

### 4.2 MA model

- Moving on to a moving average: consider  $x_t = w_t + \theta w_{t-1}$ ,  $t = 0, \pm 1, \pm 2, \pm 3, \ldots$ , where  $w_t$ ,  $t = 0, \pm 1, \pm 2, \pm 3, \ldots$  is a white noise sequence. The right-hand side here is not exactly the same as an equal-weights moving average (as we've looked at many times before), but a linear filter of the past, with weights  $a_0 = 1$  and  $a_1 = \theta$ . It is generally what we'll call a moving average (MA) model in the context of ARIMA, as we'll learn soon
- By direct calculation, which you can check, the auto-covariance function is:

$$\gamma(h) = \begin{cases} (1+\theta^2)\sigma^2 & h = 0\\ \theta\sigma^2 & |h| = 1\\ 0 & |h| > 1 \end{cases}$$

where  $\sigma^2$  denotes the variance of  $w_t$ 

• The spectral density, according to (11), is therefore

$$f(\omega) = (1+\theta^2)\sigma^2 + \theta\sigma^2(e^{2\pi i\omega} + e^{-2\pi i\omega})$$
$$= \sigma^2 \Big(1+\theta^2 + 2\theta\cos(2\pi\omega)\Big)$$

where in the second line we used  $\cos(\theta) = (e^{i\theta} + e^{-i\theta})/2$ , which follows from Euler's formula  $e^{i\theta} = \cos(\theta) + i\sin(\theta)$ 

• So an MA process has a spectral density that decays from zero, and larger  $\theta$  means a steeper decay from  $\omega = 0$  to  $\omega = 1/2$ . See Figure 6 for an illustration

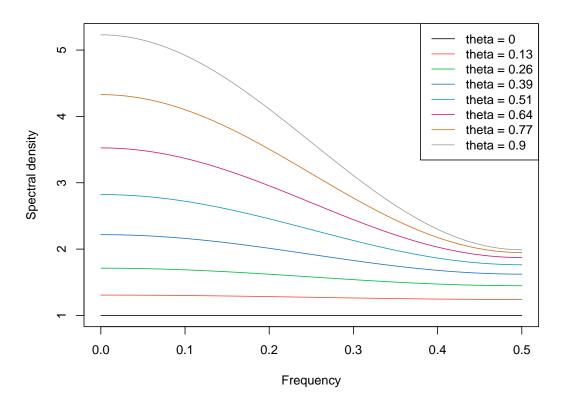


Figure 6: Spectral density for the MA model, as  $\theta$  varies.

### 4.3 AR model

- Lastly, we'll turn to an autoregressive model: consider  $x_t = \phi_1 x_{t-1} + \phi_2 x_{t-2} + w_t$ ,  $t = 0, \pm 1, \pm 2, \pm 3, \ldots$ , where  $w_t, t = 0, \pm 1, \pm 2, \pm 3, \ldots$  is a white noise sequence. This is a second-order autoregressive (AR) model, which we'll learn more about when we cover ARIMA soon
- Calculating the auto-covariance function is going to a bit nasty for this process (as you'll learn more about later when we cover ARIMA). However, for our purposes here, we can get away with a trick: finding an equation that relates the auto-covariance of white noise to that of the AR process
- Since  $w_t = x_t \phi_1 x_{t-1} \phi_2 x_{t-2}, t = 0, \pm 1, \pm 2, \pm 3, \dots$ , we have

$$\begin{aligned} \gamma_w(h) &= \operatorname{Cov}(w_{t+h}, w_t) \\ &= \operatorname{Cov}(x_{t+h} - \phi_1 x_{t+h-1} - \phi_2 x_{t+h-2}, x_t - \phi_1 x_{t-1} - \phi_2 x_{t-2}) \\ &= \gamma_x(h) - \phi_1 \gamma_x(h+1) - \phi_2 \gamma_x(h+2) - \phi_1 \gamma_x(h-1) + \phi_1^2 \gamma_x(h) + \phi_1 \phi_2 \gamma_x(h+1) \\ &- \phi_2 \gamma_x(h-2) + \phi_1 \phi_2 \gamma_x(h-1) + \phi_2^2 \gamma_x(h) \\ &= (1 + \phi_1^2 + \phi_2^2) \gamma_x(h) - \phi_1 (1 - \phi_2) \big( \gamma_x(h-1) + \gamma_x(h+1) \big) - \phi_2 \big( \gamma_x(h-2) + \gamma_x(h+2) \big) \end{aligned}$$

• Now represent the auto-covariance function  $\gamma_x$  as an integral with respect to the spectral density  $f_x$ :

$$\gamma_w(h) = \int_{-1/2}^{1/2} \left( (1 + \phi_1^2 + \phi_2^2) - \phi_1 (1 - \phi_2) (e^{-2\pi i\omega} + e^{2\pi i\omega}) - \phi_2 (e^{-4\pi i\omega} + e^{4\pi i\omega}) \right) \exp(2\pi i\omega h) f_x(\omega) \, d\omega$$

• However, note that the representation of  $\gamma_w$  in terms of its own spectral density  $f_w$  is of course

$$\gamma_w(h) = \int_{-1/2}^{1/2} \exp(2\pi i\omega h) f_w(\omega) \, d\omega$$

• Since Fourier transforms are uniquely identifying, we infer that the integrands in the last two display must match:

$$\left((1+\phi_1^2+\phi_2^2)-\phi_1(1-\phi_2)(e^{-2\pi i\omega}+e^{2\pi i\omega})-\phi_2(e^{-4\pi i\omega}+e^{4\pi i\omega})\right)\exp(2\pi i\omega h)\,f_x(\omega)=f_w(\omega)$$

• And based on our earlier calculation for white noise, we know that  $f_w(\omega) = \sigma^2$ , the noise variance, so plugging that into the above, we learn

$$f_x(\omega) = \frac{\sigma^2}{(1+\phi_1^2+\phi_2^2) - \phi_1(1-\phi_2)(e^{-2\pi i\omega} + e^{2\pi i\omega}) - \phi_2(e^{-4\pi i\omega} + e^{4\pi i\omega})}$$
$$= \frac{\sigma^2}{(1+\phi_1^2+\phi_2^2) - 2\phi_1(1-\phi_2)\cos(2\pi\omega) - 2\phi_2\cos(4\pi\omega)}$$

where in the second line we used the identity  $\cos(\theta) = (e^{i\theta} + e^{-i\theta})/2$  once again

• Plotting this, as we do in Figure 7, we learn that a second-order AR process has a spectral density that is concentrated around a particular frequency. For example, when  $\phi_1 = 1$  and  $\phi_2 = -0.9$ , it is concentrated around  $\omega \approx 0.16$  (a cycle of  $1/\omega \approx 6$  time points)

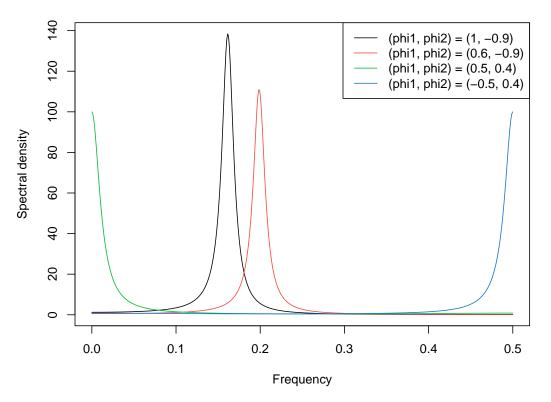


Figure 7: Spectral density for the second-order AR model, for a few parameter choices  $\phi_1, \phi_2$ .

#### 4.4 Sample spectral density

- To conclude, let's draw a connection between the two main players that we've seen in this lecture: the periodogram, and the spectral density
- Recall the DFT (9) of a series  $x_t, t = 1, ..., n$ , which we can rewrite as

$$d(\omega_j) = \frac{1}{\sqrt{n}} \sum_{t=1}^n x_t \exp(-2\pi i \omega_j t)$$

for  $\omega_j = j/n, \, j = 0, 1, \dots, n-1$ 

• We claim that  $\sum_{t=1}^{n} \exp(-2\pi i\omega_j t) = 0$  for any j < n. This can be check by viewing it as a geometric series  $\sum_{t=1}^{n} z^t$  in  $z = e^{2\pi i j/n}$ , and hence we can use the formula

$$\sum_{t=1}^{n} z^{t} = z \frac{1-z^{n}}{1-z}, \quad \text{for } z \neq 1$$

but here  $z^n = e^{2\pi i j} = 1$  for any j < n

• Using the last fact, we can thus rewrite the DFT once more as

$$d(\omega_j) = \frac{1}{\sqrt{n}} \sum_{t=1}^n (x_t - \bar{x}) \exp(-2\pi i \omega_j t)$$

• The periodogram, recalling (10), is given by the squared modulus of the DFT,  $P(\omega_j) = |d(\omega_j)|^2$ . Using the fact that the squared modulus of z = a + bi is  $|z|^2 = a^2 + b^2 = (a + ib)(a - ib)$ ,

$$P(\omega_j) = \frac{1}{n} \sum_{s=1}^n \sum_{t=1}^n (x_s - \bar{x})(x_t - \bar{x}) \exp\left(-2\pi i\omega_j(t-s)\right)$$
  
=  $\frac{1}{n} \sum_{h=-(n-1)}^{n-1} \sum_{t=s+h} (x_s - \bar{x})(x_{s+h} - \bar{x}) \exp(-2\pi i\omega_j h)$   
=  $\sum_{h=-(n-1)}^{n-1} \frac{1}{n} \sum_{s=1}^{n-|h|} (x_s - \bar{x})(x_{s+|h|} - \bar{x}) \exp(-2\pi i\omega_j h)$   
=  $\sum_{h=-(n-1)}^{n-1} \hat{\gamma}(h) \exp(-2\pi i\omega_j h)$ 

where  $\hat{\gamma}(h)$  is our usual sample auto-covariance at lag h

- Compare the last line to the spectral density (11), and what do you find? The periodogram is simply a sample spectral density estimator, at discrete frequencies  $\omega_j$ ,  $j = 0, \ldots, n-1$
- For models in which we expect smoothness in the spectral density, the sample spectral density (periodogram) turns out to be not such a good estimator. There is much more we can do, and much more to say about spectral analysis and filtering in general. See Chapters 4.4–4.10 of the SS book if you're interested in learning more